



Gakugei 東京学芸大学  
Tokyo Gakugei University



# Practical nano-satellite monitoring of ionospheric earthquake precursors

Masashi Kamogawa, Hide Nitta, Shoho Togo

Department of Physics, Tokyo Gakugei University, Tokyo, Japan

Toshiyasu Nagao

Division of Earthquake Prediction, Volcano and Tsunami Researches

Tokai University, Japan

# Earthquake prediction

- Short-term earthquake prediction is still very difficult.
- For this prediction, observable precursors are required.
- Many kinds of precursors have been reported, but controversial.
- Ionospheric precursors are plausible phenomena for the prediction, because the statistical verification can simply be done using satellite.

# Plausible atmospheric ionospheric precursors

Possibly magnetic-storm origin?  
Rishbeth, EOS, 2006

F layer  
Ionosphere

Liu et al.  
JGR (2006)

No reproduction  
Clilverd et al., JGR, 1996  
Optimistic statistics  
Michael, GRL, 1997

Es layer  
E layer

Nemec et al.  
GRL (2008), JGR (2009)

??

Hayakawa et al.  
JGR (2010)

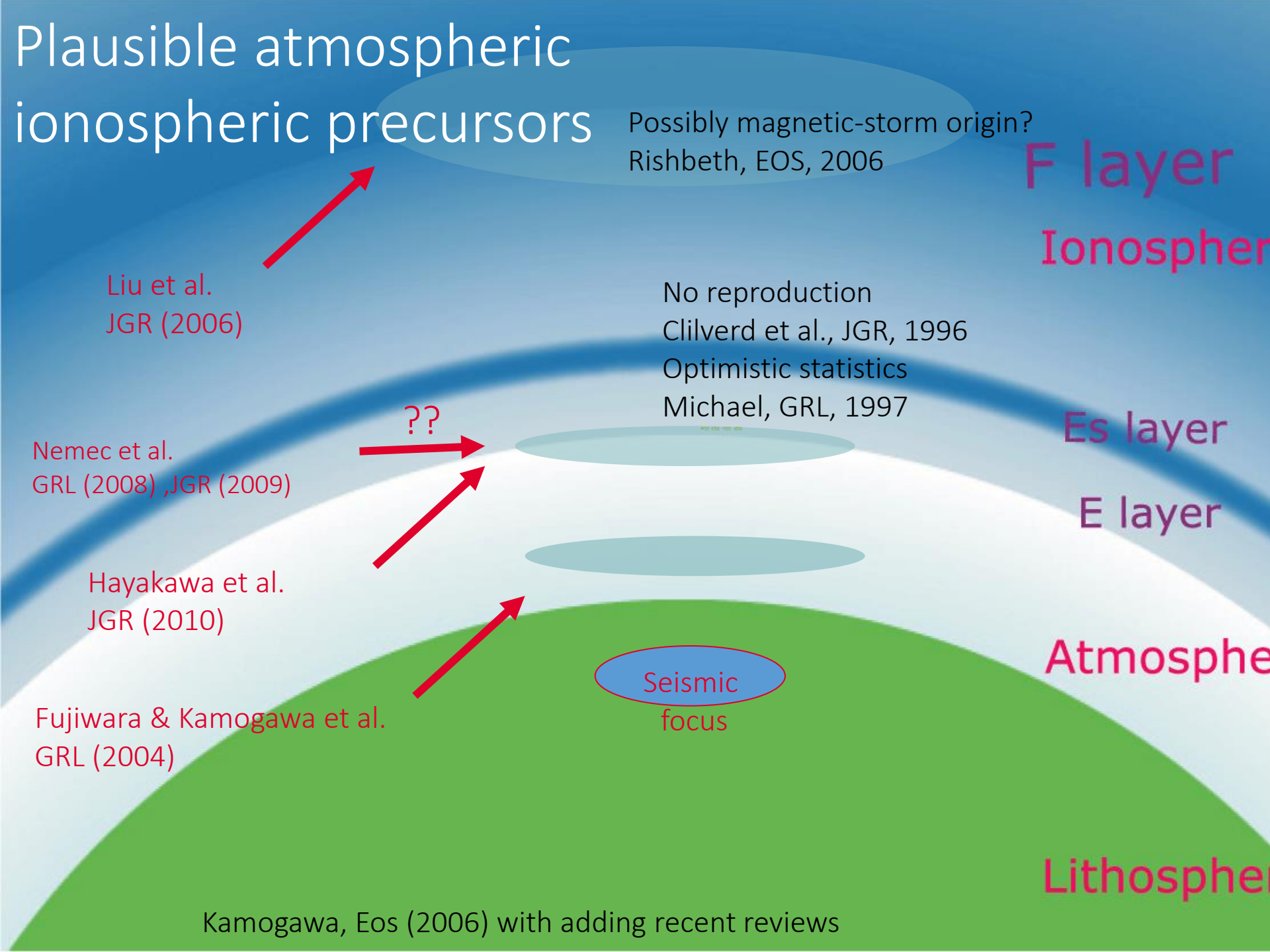
Atmosphere

Fujiwara & Kamogawa et al.  
GRL (2004)

Seismic focus

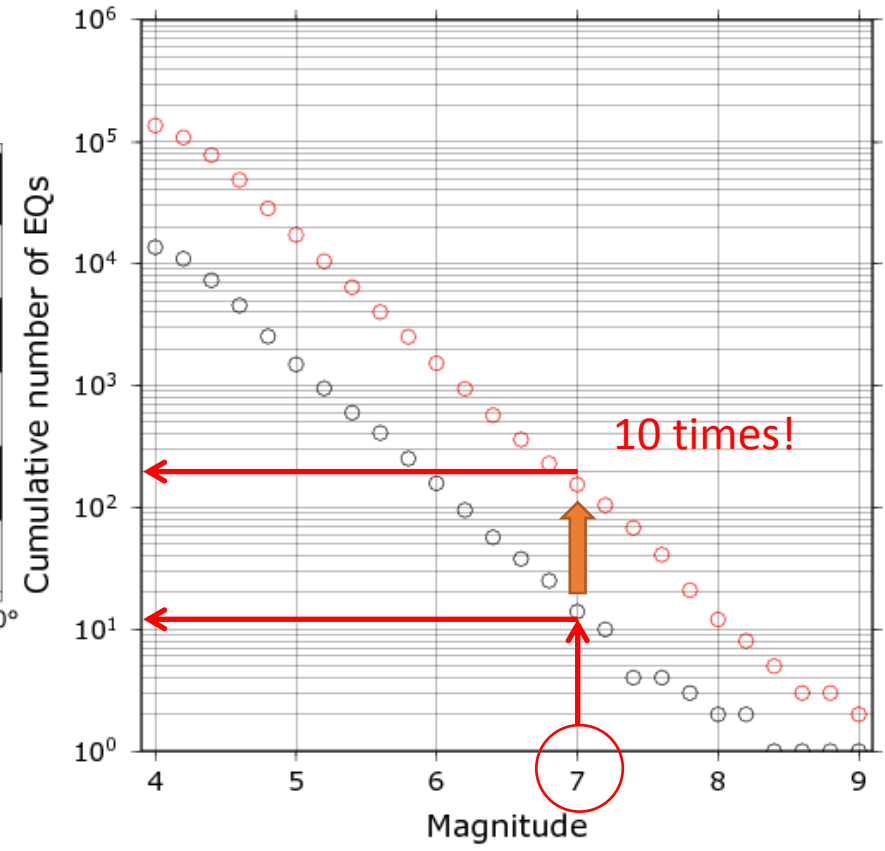
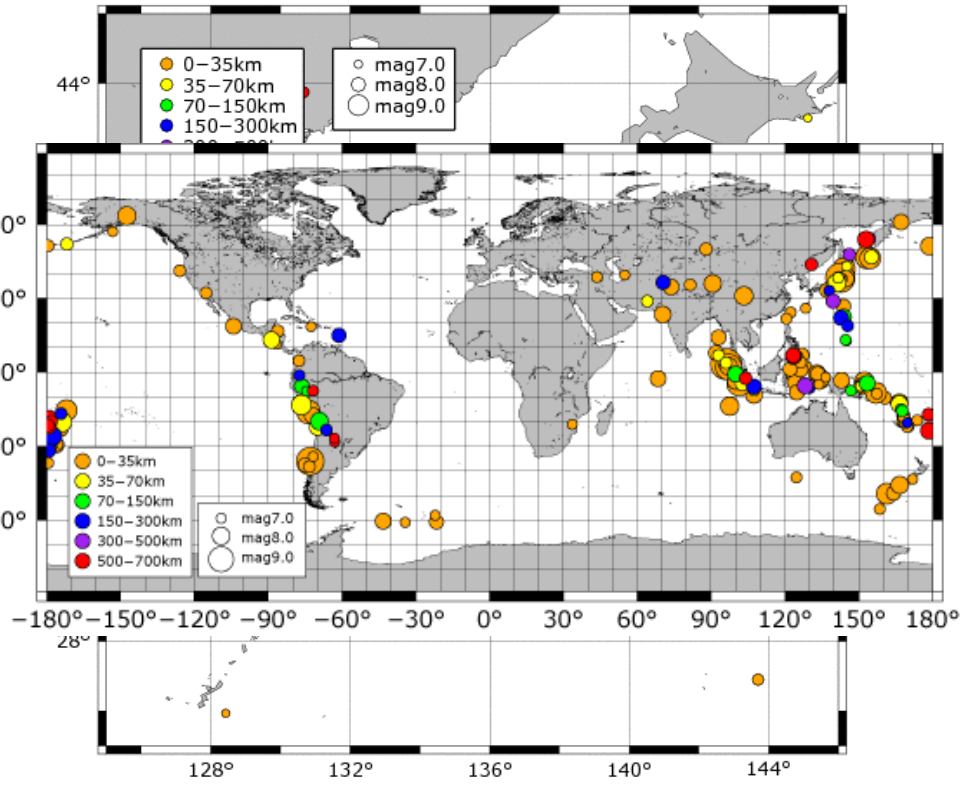
Lithosphere

Kamogawa, Eos (2006) with adding recent reviews



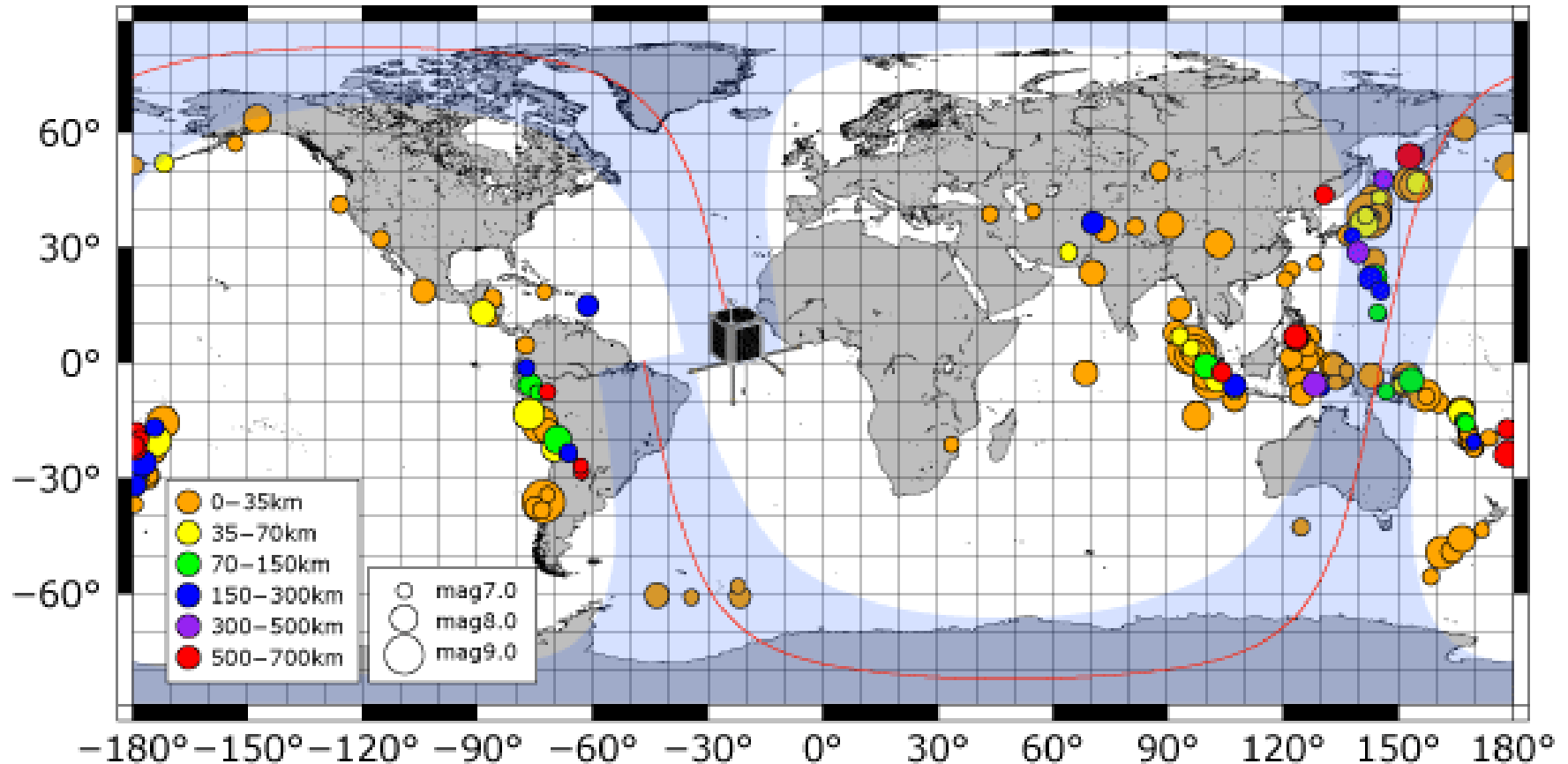
# Global seismicity

(USGS, 2000-2011)



Satellites easily monitor global seismicity!

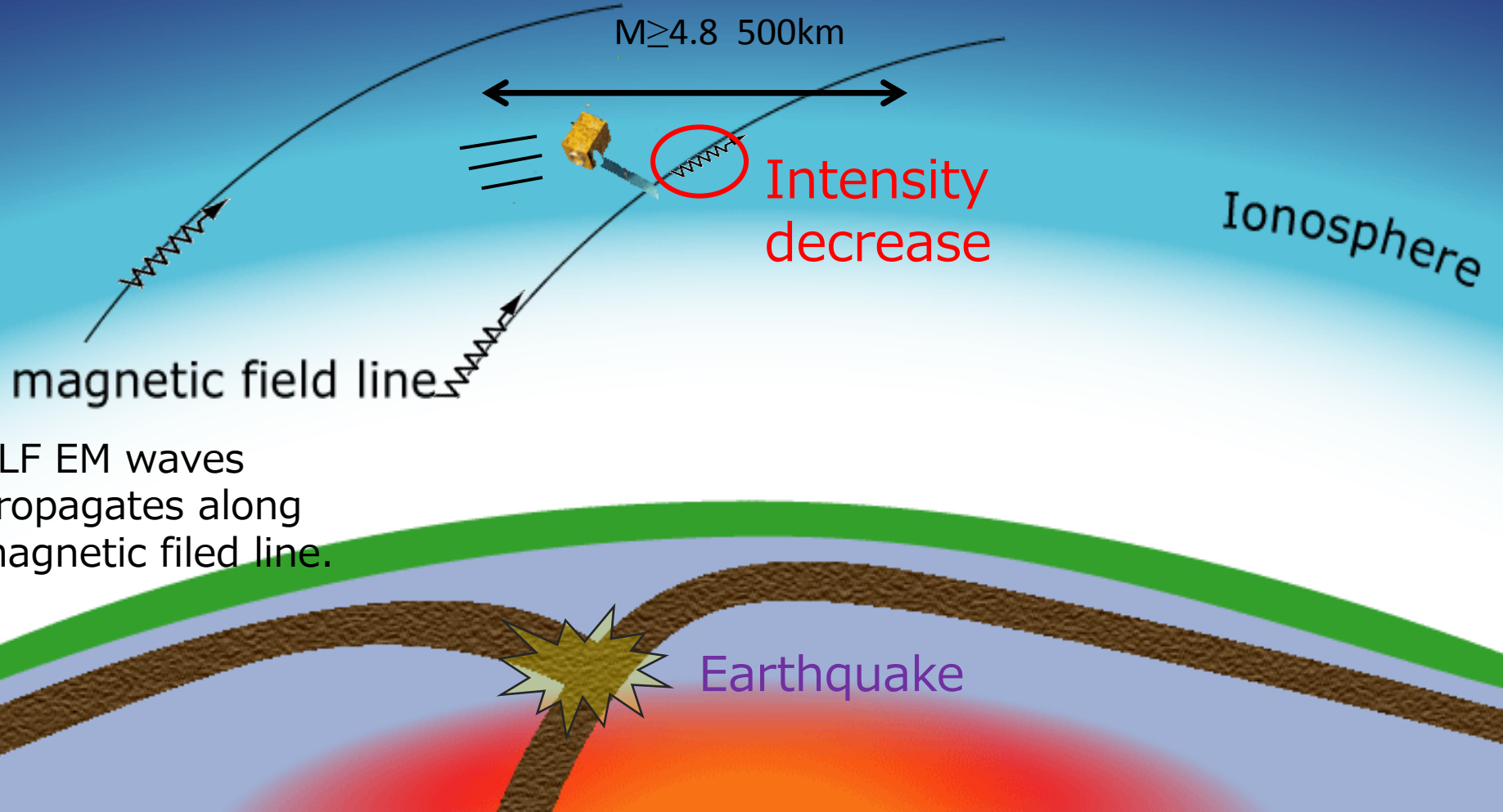
# Satellite observation is useful!



Satellites easily monitor global seismicity!

# Plausible precursor: Intensity decrease of VLF electromagnetic waves

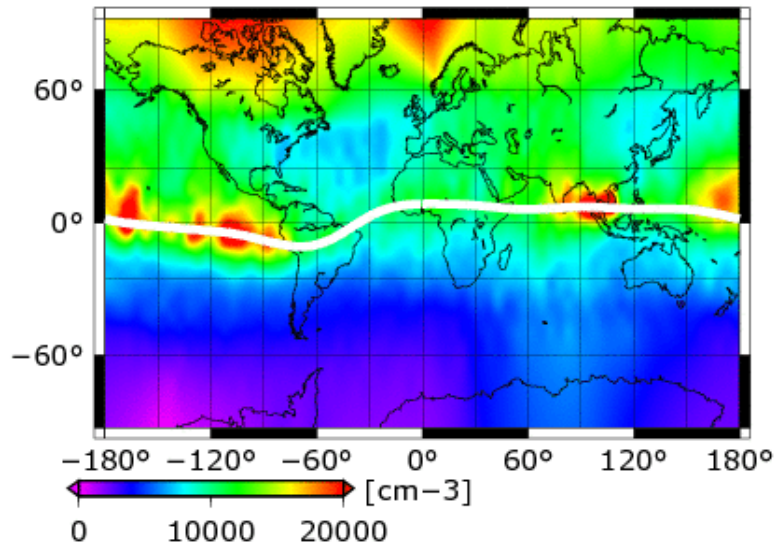
Nemec et al. (2008; 2009)



# How to identify a precursor?

We should know standard ionosphere .

So, we construct standard ionospheric model.



Example of electron density map during geomagnetically quiet period

Model depends on..

Local time

Latitude

Longitude

Altitude

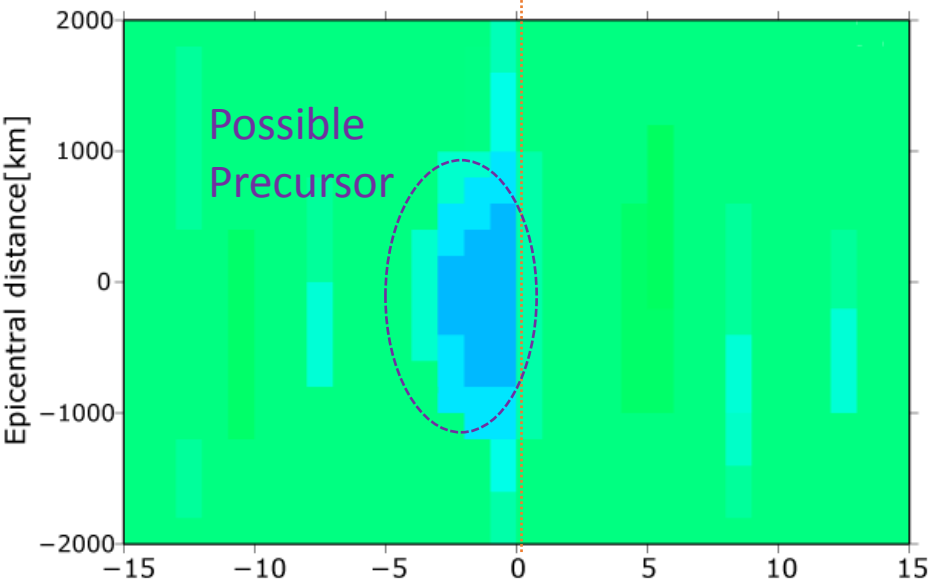
Solar flux

Geomagnetic activity

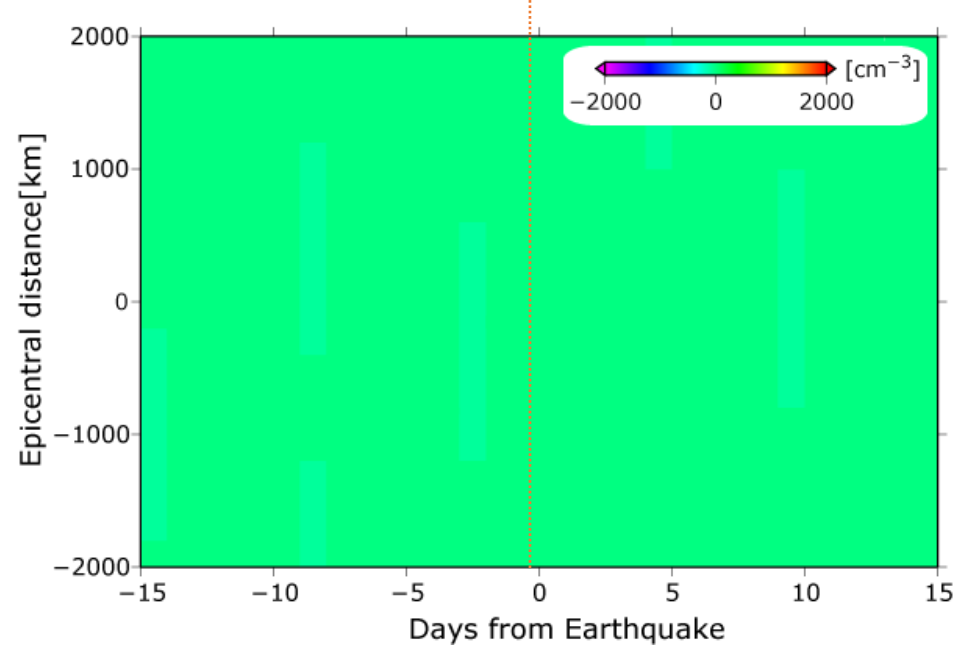
etc....

# Accumulate residual values between model and observed data.

Time of **Real** earthquakes



Time of **Virtual** earthquakes



Note: This is a conceptual view.

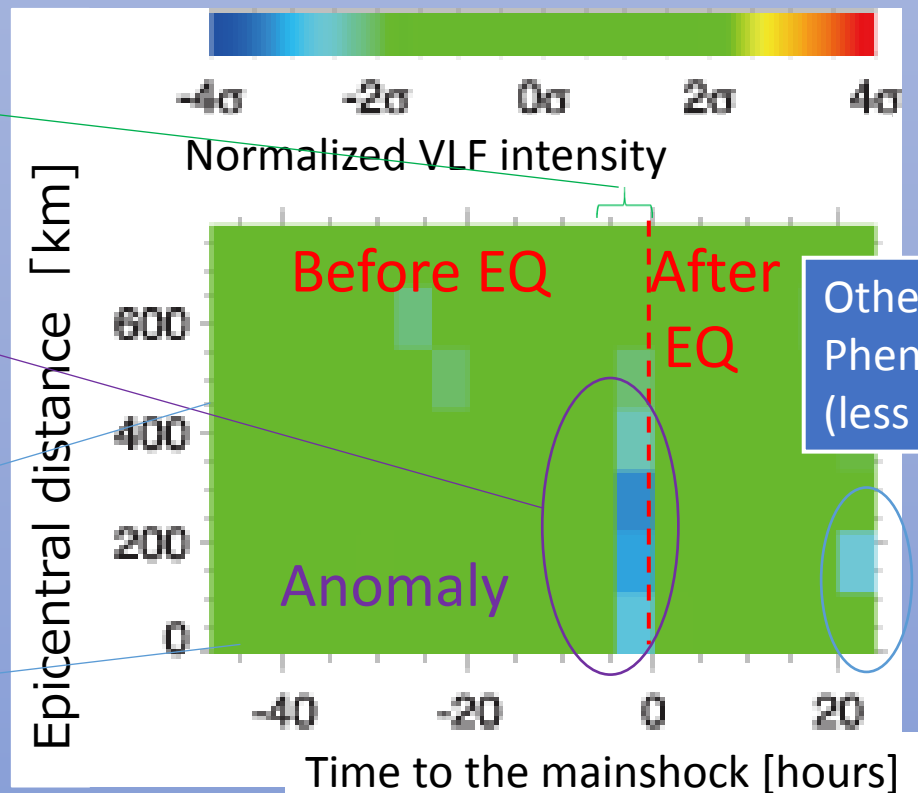


# Statistical study of pre-seismic VLF intensity decrease (=Background VLF intensity)

Superposed epoch analysis from stacking 100 events

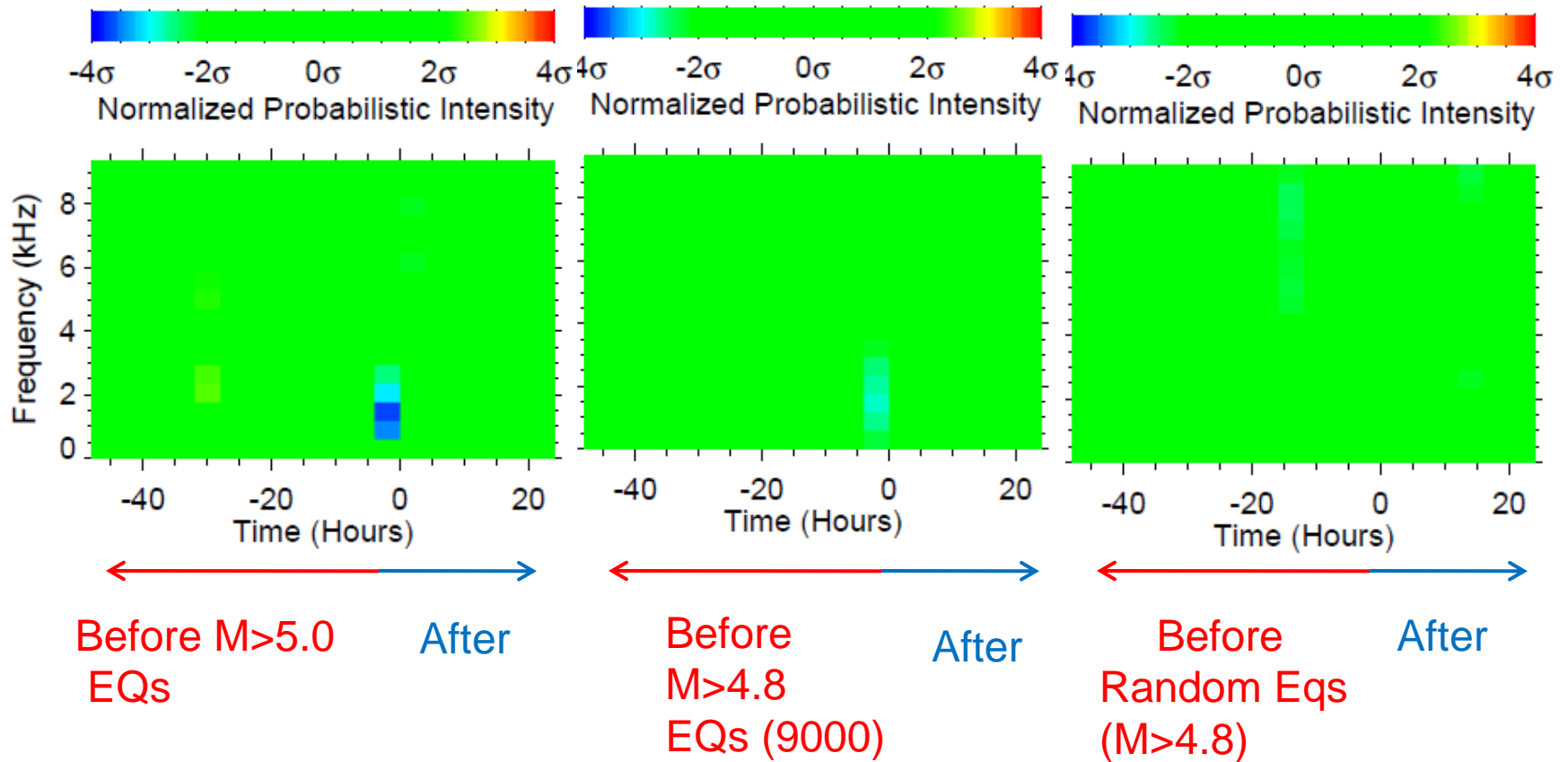
Model parameter

- M>4.8  
Within 4 hours
- More than  $3\sigma$  of normalized intensity
- Epicentral distance Within 500 km

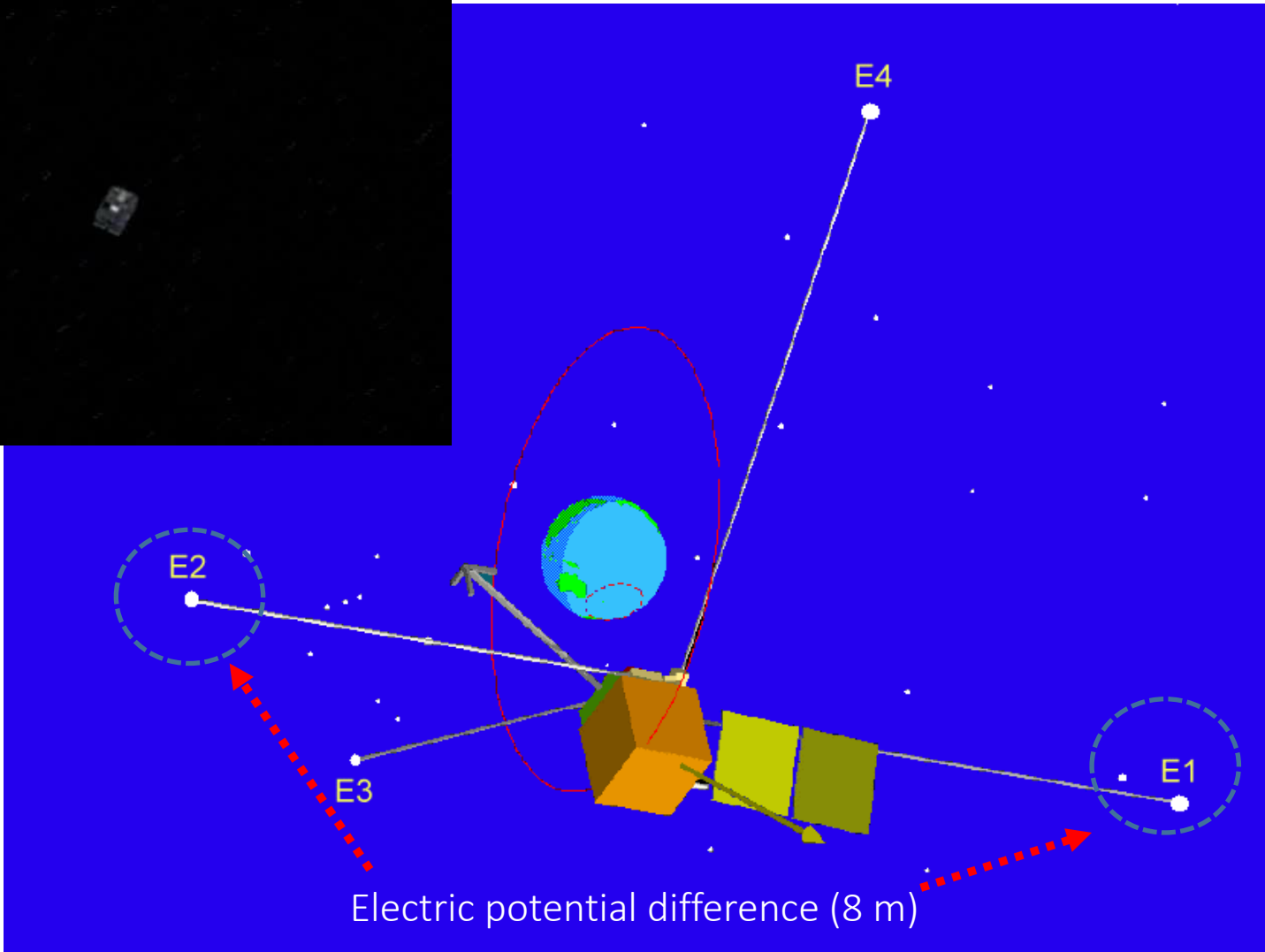


# Comparison: Large and Small Magnitude EQ

## Real and virtual EQs



# DEMETER electric field measurement



Electromagnetic wave  
observed by space craft

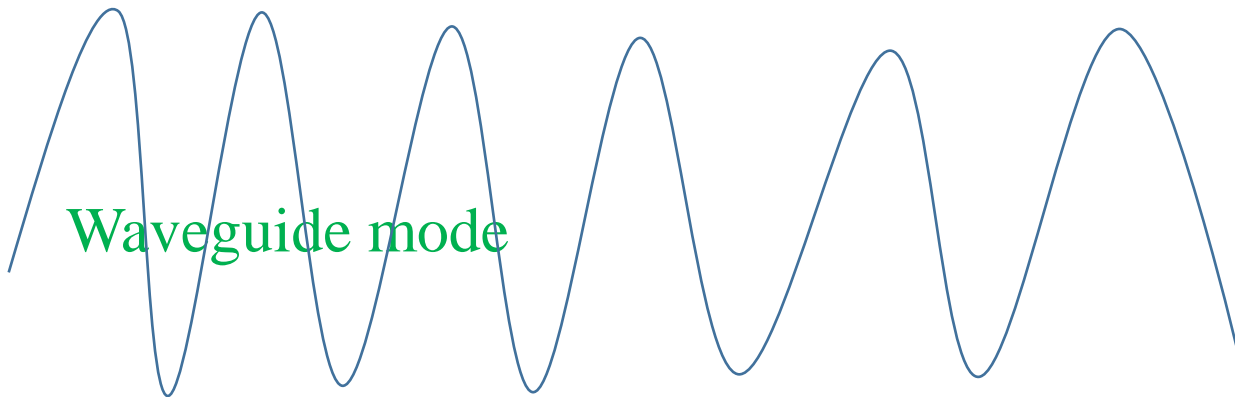
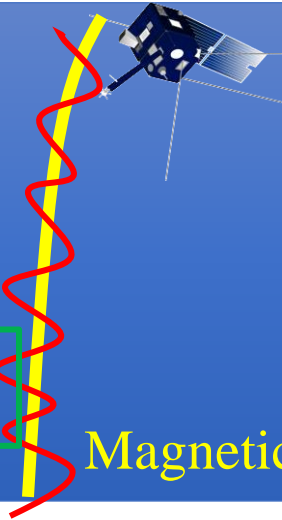
ionosphere

**Whistler-wave**

Magnetic field line

**Hiss , Chorus**  
(Magnetosphere origin)

Something noise etc..



Waveguide mode

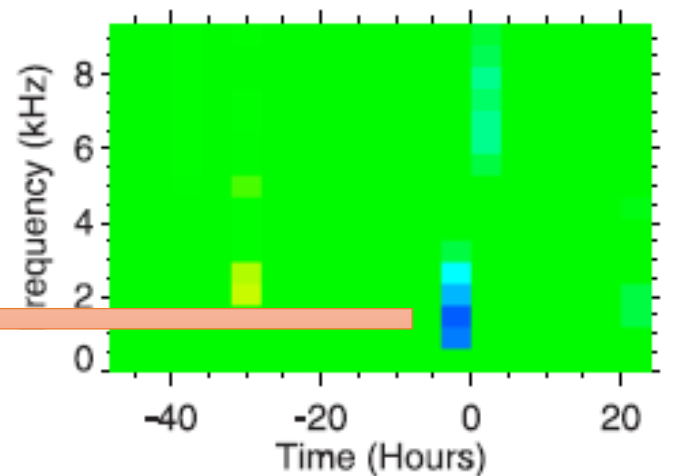
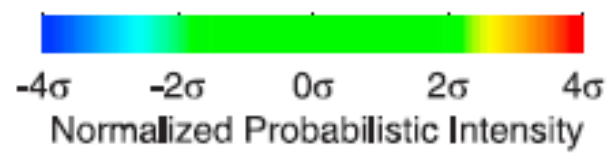
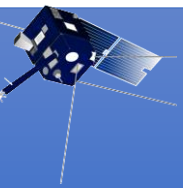
90km

ground

**Whistler-wave**

ionosphere

Magnetic field line



1.7kHz

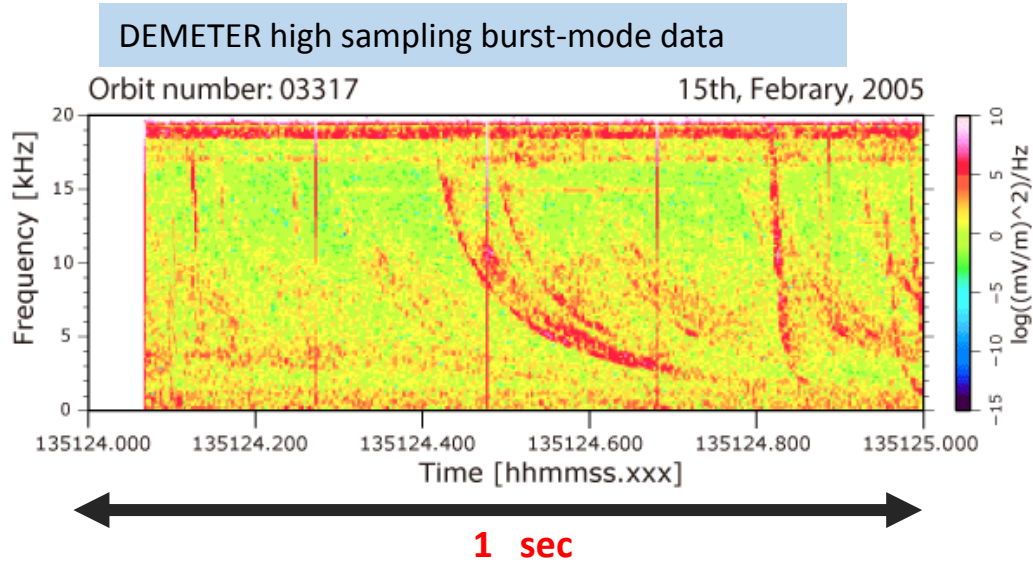
Waveguide mode



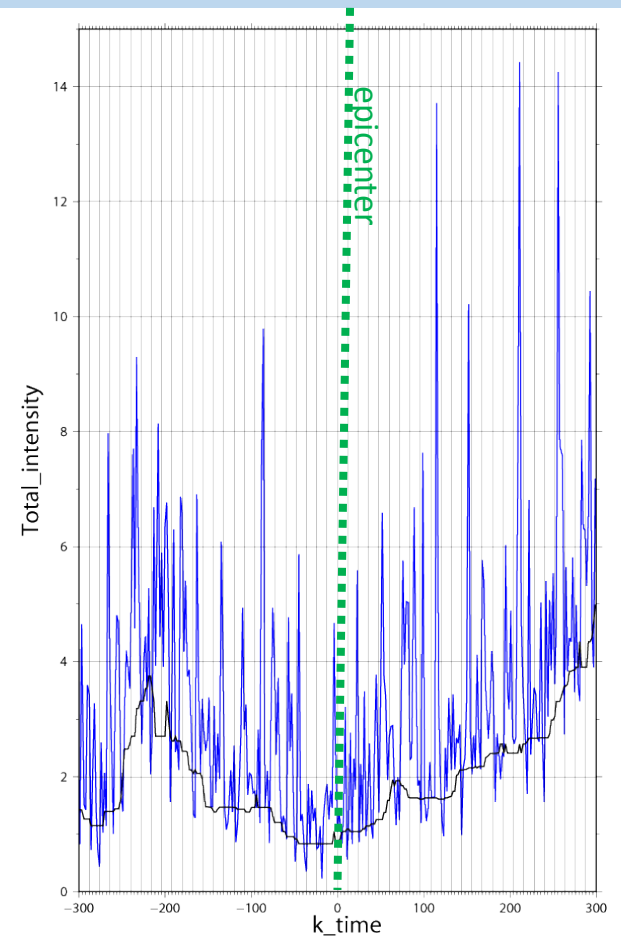
**earthquake**

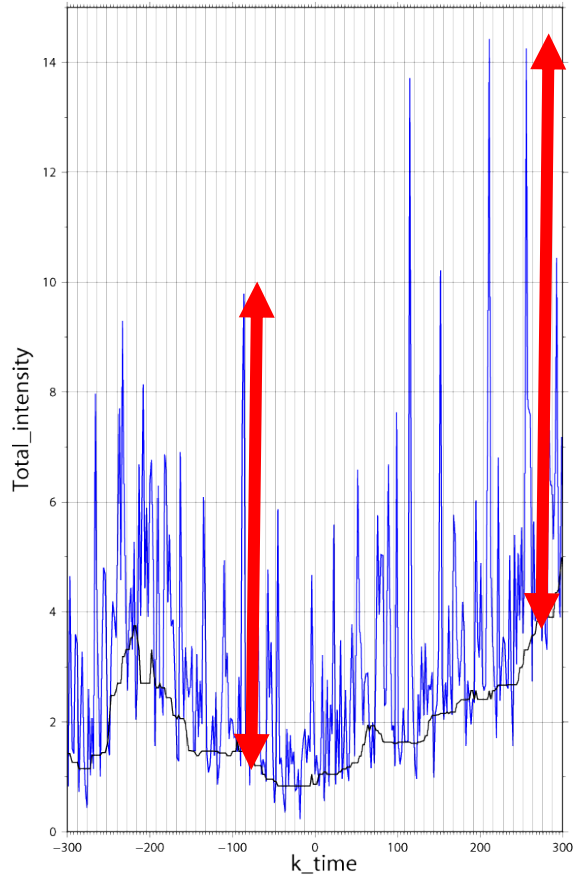
ground

# Whistler wave intensity



DEMETER 2sec sampling survey-mode data





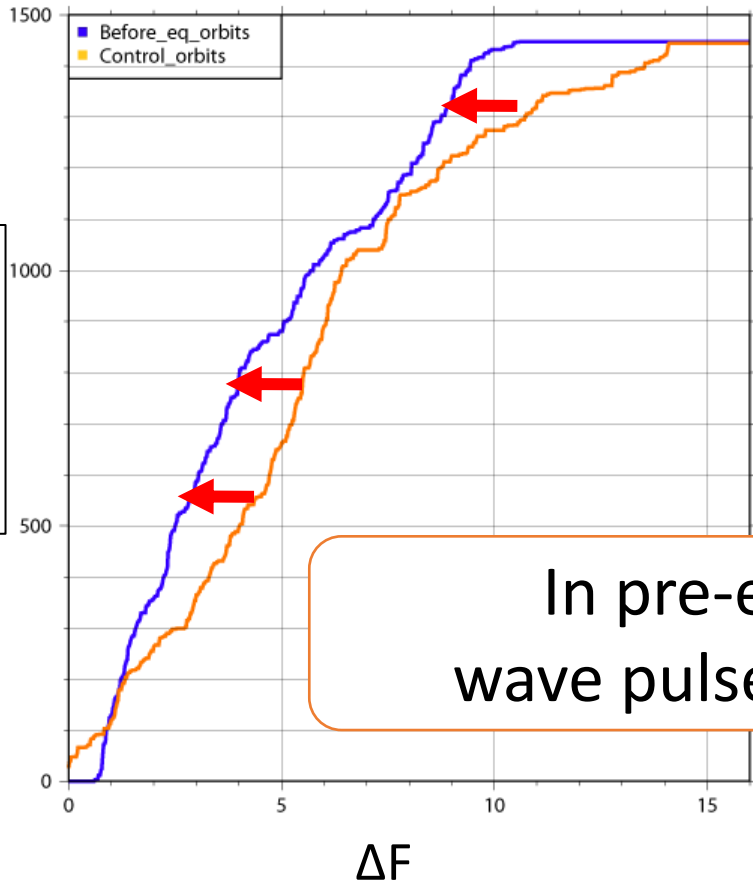
$$\Delta F = F - B$$

F : total power

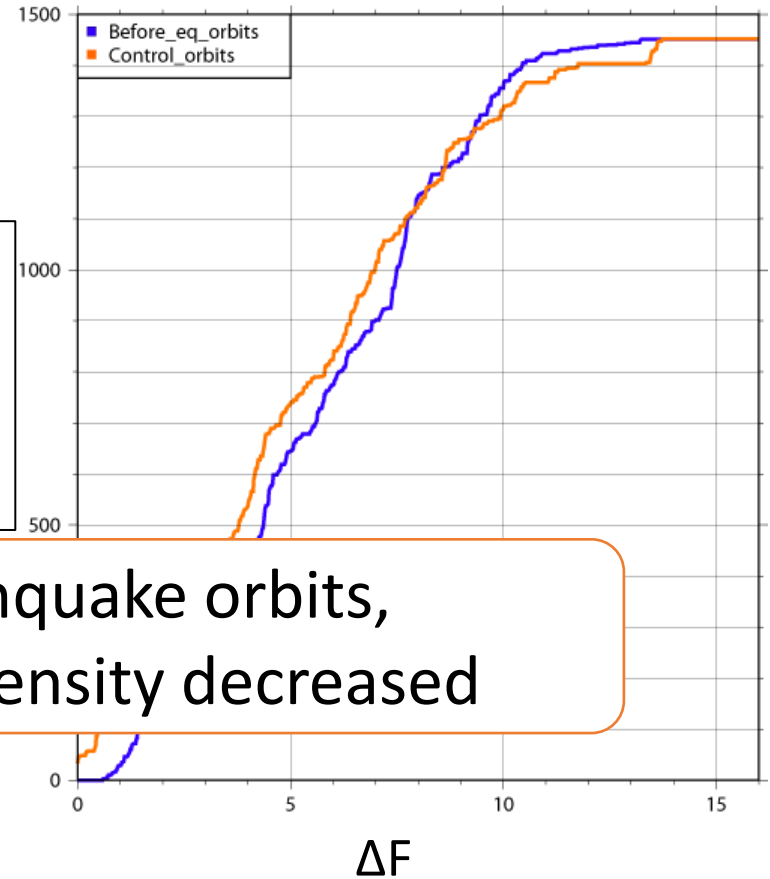
B : baseline

$\Delta F$  roughly denotes the whistler wave intensity attributed to lightning activity.

## Above the epicenter



## Far from the epicenter



In pre-earthquake orbits,  
wave pulse intensity decreased



# Conclusion: Feasible Nano-Sat

- Two spherical electrode probes (3 cm radius) are equipped with the end of tow 40 cm booms.
- Bootstrap are set up at both the sides of probe.
- Only nighttime observation
- Two mode sampling
  - a. SURVEY MODE  
(Continuous 1 sec)  
Averaged electric field of 2-8 kHz after FFT analysis of 20 kHz sampling is observed. (180KB/day)
  - b. BURST MODE  
(Event command request)  
20kHz sampling electric field for two days is overwritten in on-board memory. 300s raw data within 4 hours before the earthquake is transmitted by the command. (1.2 MB/day, 10 days = 1 EQs)

